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Project Report

ETS-12

E. W. Rork

Detections of Faint Geosynchronous
Satellites at the GEODSS ETS
from October 1976 to March 1977

10 June 1977

Prepared for the Department of the Air Force
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MASSACHUSETTS INSTITUTE OF TECHNOLOGY

LEXINGTON, MASSACHUSETTS



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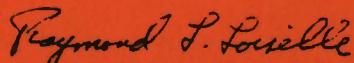
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FOR THE COMMANDER

A handwritten signature in dark ink, reading "Raymond L. Loiselle". The signature is written in a cursive style with a large, stylized initial 'R'.

Raymond L. Loiselle, Lt. Col., USAF
Chief, ESD Lincoln Laboratory Project Office

MASSACHUSETTS INSTITUTE OF TECHNOLOGY
LINCOLN LABORATORY

DETECTIONS OF FAINT GEOSYNCHRONOUS
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FROM OCTOBER 1976 TO MARCH 1977

E. W. RORK

Group 94

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ABSTRACT

This report presents a record of faint geosynchronous satellites detected at the GEODSS ETS from October 1976 to March 1977. The satellites reported have at one time or another been difficult or impossible to detect electro-optically because solar reflections from them were faint.

DETECTIONS OF FAINT GEOSYNCHRONOUS SATELLITES AT THE GEODSS ETS
FROM OCTOBER, 1976 TO MARCH, 1977.

I. Introduction. Certain geosynchronous satellites have only occasionally been detected at the GEODSS ETS because in general, solar reflections from them are too faint. Satellites as faint as $17.9 m_v$ have been detected at the ETS using TV-integration. The ETS has been searching for and detecting synchronous satellites on a regular basis since it began operations in September 1975, partly to discover any difficulties in detecting them electro-optically as a function of night-sky brightness, atmospheric extinction, the sun-observer-satellite angle, and physical characteristics of the satellite itself. For example, certain synchronous satellites were not detected during the summer solstice, 1976, but were detected at other times of the year.¹ At the time of a summer solstice, the high declination of the sun causes specular reflections from synchronous satellites with surfaces oriented parallel to the earth's axis to be reflected southward, hence such satellites must be detected from diffuse solar reflections.²

The ETS now has a greater detection capability by the use of an automatic moving-target-indicating (MTI) device, and TV-integration. This report summarizes observations of a series of synchronous satellites at the ETS which have occasionally not been detectable. Since reasonable attempts were made to eliminate

telescope pointing as the reason for not detecting a satellite, the most likely reason was that it was too faint. The observations reported are part of a project now being carried out at the ETS to determine when and if, during the year, certain synchronous satellites are not detectable.

II. Current Detection Capability.

Three methods of confirming the presence of a satellite in a field of view (FOV) are currently used at the ETS:

1. Normal TV-Rate Display. After the computer-controlled telescope has been directed to a computed satellite trajectory, the operators may first try to spot the satellite moving with respect to stars displayed on a video monitor. Using the 31-inch ETS sensor in a clear dark sky, a satellite as faint as about 16.5 visual magnitudes can be spotted in the 1° FOV, and a satellite as faint as about 17.5 visual magnitudes can be spotted in the 0.5° FOV. The visual magnitude figures are given for the case of no atmospheric extinction. Such detections, however, require the operators to be very alert.

2. Automatic Moving Target Indicating Device. A device has been recently installed at the ETS which provides:

1. a video display with star images cancelled out, leaving on the display any object which moves from one resolution cell to another during a preset time interval, and
2. an indication alarm and mark around the object on the TV monitor if it is present.

The automatic detection sensitivity achieved is about 15.5-16.0 visual magnitudes in the 1° FOV, and 16.3-16.9 visual magnitudes in the 0.5° FOV. However, operators must

visually inspect the display to confirm the presence of a moving satellite, since a few false alarms may accompany a faint detection. In addition, the operators may visually detect satellites as faint as those able to be detected on the TV-rate display by examining the processed video display of the MTI device. Faint detections are easier, because stars have been removed from the display, and a TV-integration of up to 8 video frames improves the signal-to-noise ratio produced by the image of a faint satellite.

The usefulness of the MTI device is most apparent when a large area of the sky must be searched to find a satellite. In this case, a large number of fields must be sequentially examined in a step-stare search program.

3. TV-Integration. To search for a satellite suspected of being fainter than the 17th visual magnitude, integration is used both on the TV camera and an external storage device to increase the video signal-to-noise ratio from a possible satellite image. In such an operation, it is necessary for the image of the satellite to remain in the same resolution cell of the TV sensor during the entire integration time, which is typically about 9 seconds. Since the angular diameter of a resolution cell of the current 31-inch ETS sensor is either 5 or 10 arc seconds, corresponding to the 0.5° or 1° FOV, respectively, and since a synchronous

satellite moves at an angular velocity of about 15 arc seconds/sec eastward with respect to the stars, it is necessary to drive the telescope close enough to the angular velocity of the satellite to keep its image in a resolution cell during the integration time. If the SNR of a satellite in the field reaches 10, it can be detected at a glance against the streaking stars on a video monitor; however, if it is only about 3, it is very hard to detect visually. In the latter case, the telescope must be moved to track at an offset from the original position and the integration process repeated to see if the faint object has moved. Frequently, this procedure must be repeated several times to confirm the presence of a faint satellite. In dark, clear skies, using the 0.5° FOV of the 31-inch ETS sensor, an 18th visual magnitude object can produce a $\text{SNR} = 10$ after about 9 seconds of integration. Integrating beyond 9 seconds doesn't help in the present set-up because the video signal due to sky background and/or dark current becomes too prominent. Future experiments are planned in which the target of the camera tube will be cooled to reduce its dark current to permit longer integration times.

III The Satellites Observed.

A historical record of the observations made on the satellites is presented on Table 1. Figure 1 is a graphical record of the detections as a function of the day, sun declination, and lunar phase. A few brightness measurements have been reported in Table 1, primarily for the faintest detections.

We notice that:

1. 83501, 83502, 83512 were fainter around the winter solstice, 1976, (day 6356), when specular reflections would be furthest north.
2. 83546, 83547, 83550, the IDCSP satellites, were not detected near the time of the winter solstice, and were very faint at other times.

IV. Conclusions.

Satellites were detected as faint as 17.9 m_v in these observations. It is probable that some of the satellites in Table 1 were fainter than 17.9 m_v when they were not detected. If it is necessary to routinely detect satellites fainter than the 18th visual magnitude, it will be necessary to improve the E-O detection process for GEODSS. This can be done by increasing telescope aperture, improving detector resolution, increasing integration time, and improving MTI efficiency, possibly with considerable software treatment to the data. It will be difficult to accomplish this without increasing search time appreciably.

The observations reported are part of a project now being carried out at the ETS which is to detect the satellites listed in Table 1 on a regular basis if possible, up to and through the summer solstice, 1977. A record of such observations will indicate:

1. If and when the satellites become too faint for TV rate detection;
2. How faint they become; and
3. If and when they are too faint to be detected by integration at the ETS.

Such information, when correlated with the sun-observer-satellite angle and night-sky conditions, will help establish a necessary detection sensitivity threshold for the GEODSS program.

ACKNOWLEDGEMENTS

Many GEODSS ETS site personnel of M.I.T., Lincoln Laboratory and the U.S.A.F. Air Defense Command have participated in the searches for and detections of the satellites reported herein, and their work is greatly appreciated. The author thanks Dr. J. M. Sorvari for his estimations of the brightness of 83547 and 83550 on UTD 6290.

REFERENCES

1. E.W. Rork, "Brightness Measurements Taken With The GEODSS Experimental Test System on Synchronous Satellites During the Summer Solstice, 1976," in the Proceedings of the Eighth NORAD Spacecraft Identification Conference (NORSIC-8), 2-6 August 1976 (to be published).
2. A.S. Friedman, "Determination of Specular Reflection from Cylindrical Satellites for Electro-Optical Surveillance and SOI" Project Report ETS-3, Lincoln Laboratory, M.I.T. (8 October 1976), DDC AD-A03458011.

TABLE 1. A Record of Detections of and Searches for Faint Synchronous Satellites at the GEODSS ETS from October, 1976 to March, 1977.

Legend: B: Satellite was bright enough to be readily detected.
F: Satellite was faint, and may have been hard to bore-sight. In general, the symbol F was used when the satellite visual magnitude was fainter than about 15.5.

The standard deviation in the brightness measurements is estimated to be $0.2 m_v$ except where a \sim sign appears, in which case it could be 0.3 or $0.4 m_v$.

TABLE 1.a

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO. OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83501	NATO 2	6323		1			
		6328		1			
		6329	2		X	~17	
		6330	1	1	X	17.0	
		6338		1			
		6342		1			
		6356	2		X	17.1	
		6357	2		X	~17	
		6358	1		X	17.0	
		6363	1	3	X	17.1	
		6364	2	5	X	16.4	
		7007		1	X		
		7011		1			
		7014		1			
		7025	1			F	
		7028	3		X	F	
		7029		2			
		7039		1			
		7040	1			15.3	
		7045	2			15.4, 16.0	
		7048	1			B	
		7049	2	1		F	
		7050	2		X	F	
		7063	4			B	
		7067	1			B	
		7068	1			B	
		7069	1			F	
		7073	1			B	
		7074	1		X	F	
		7075	2			B	
		7077	1			F	

TABLE 1.b

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83502	NATO 1	7078	1			B	
		7082	1			F	
		7083	1			B	
		7089	1			F	
		6287		1			
		6294		1			
		6304		1			
		6323		1			
		6328		2			
		6329	1		X	~17	
		6330	1		X	17.0	
		6338		1			
		6339		2			
		6349		1			
		6351		1			
		6356	3		X	17.1	
		6357	2		X	~17	
		6358	1		X	17.0	
		6363	1	3	X	17.1	
		6364	1	2	X	17.4	
		7011		1			
		7012		1			
		7014		1			
		7020		1	X		
		7025	1	1		F	
		7028	2		X	F	
		7029		1			
		7040	1			15.3	
		7045	2			15.5	

TABLE 1.c

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83503	Skynet A	7048	2			B	
		7049	3			F	
		7050	2			F	
		7061	1			B	
		7063	1			B	
		7068	1			B	
		7069	1			F	
		7073	1			B	
		7074	1		X	F	
		7075	2			B	
		7077	1			F	
		7078	1			B	
		7082	1	1		B	
		7083	2			F	
		7089		2			
		6304		1			
		6328		2			
		6329		1	X		
		6330		1	X		
		6337	1			B	
		6338	1			B	
		6339		2			
		6342		1			
		6346		1			
		6356		3	X		
		6357		3	X		
		6358		1	X		
		6364		4	X		
		7025		1			
		7048	2			B	

TABLE 1.d

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83505	SMS-3	7049	2			F	
		7050	2			B	
		7063		1			
		7074		1			
		7077		1			
		7078		1			
		7079		1			
		6272	2			B	
		6273	2			B	
		6275	2			B	
		6279	1			B	
		6280	8			B	
		6281	3			B	
		6283	4			B	
		6294	1			B	
		6316	1			B	
		6317		1			
		6346	3			F	
		6349	1	2		B	
		6251	1			B	
		6364		1	X		
		7004		3			
		7007		2	X		
		7012	1			B	
		7013	2			B	
		7014	1			B	
		7019	8	1		B	
		7028	1			B	
		7029		2			
		7035		1			

TABLE 1.e

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83507	IS II F-3	7048	1			B	
		7049	2			F	
		7063	4			B	
		7064	4			B	
		7065	2			B	
		7066	3			B	
		7067	3			B	
		7068	3			B	
		7071	2			B	
		7073	1			B	
		7078	1			B	
		7089	1			B	
		6279		1			
		6287		1			
		6304		1			
		6314		1			
		6317	1			B	
		6330		3	X		
		6339		1			
		6356		2	X		
		6358		2	X		
		7012	1			B	
		7019		2			
		7028		1			
		7045	1			16.3	
		7048	1			B	
		7049	3			B	
		7050	2			F	
		7062		1			
		7067		1			
		7069	1			F	

TABLE 1.f

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83509	Westar 1	7071	1			B	
		7074	2			B	
		7075	1			B	
		7077	1			B	
		7082	1			B	
		7083	2			B	
		7089	1			B	
		6281		1			
		6364		1	X		
		7011	1			B	
		7012	1			B	
		7013	1	1		B	
		7019		1			
		7020	2			B	
		7028	1			B	
		7035	1			B	
		7048	1			B	
		7049	1			B	
		7063	3			B	
		7064	1			B	
		7065	3			B	
		7066	3			B	
		7067	2			B	
		7068	1			B	
		7075	1			B	
		7083	1			B	
		7089	1	1		B	
83512	IS I F-1	6287		1			
		6294		1			
		6304		1			

TABLE 1.g

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO. OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83513	IS III F-2	6363	2	2	X	17.9	
		6364	1	2	X	17.0	
		7028		1			
		7045	2		X	15.2, 16.9	
		7048		3			
		7049		2			
		7050		2			
		7063		1			
		7071		1			
		7073	1			B	
		7074		2			
		7078		1			
		7082		1			
		7083		1			
		7089		2			
		6272	1			B	
		6273	2	2		B	
		6274	4	2		B	
		6275	1			B	
		6279	1			B	
		6280	8			B	
		6281	2			B	
		6282	2	2		B	
		6283	3	2		B	
		6287	1			B	
		6294		1			
		6295	1			B	
		6304		1			
		6307		2			
		6314		1			
		6363		3	X		

TABLE 1.h

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83523	ATS-5	7012		1			
		7014	1			F	
		7019		2			
		7045	1			15.9	
		7048	1			B	
		7049	1			B	
		7050	2			F	
		7067	2			B	
		7063	5	1		B	
		7064	2			B	
		7065	2			B	
		7066	2			B	
		7067	1			B	
		7068	1			B	
		7069	1			F	
		7073	2			F	
		7074	1			B	
		7075	1			B	
		7077	1			B	
		7078	1			B	
		7082	1			B	
		7083	2			F	
		7089	1			F	
		6272	2			B	
		6263	4			B	
		7274	4	1		B	
		6275	3			B	
		6279	2			B	
		6280	1			B	
		6294	1			B	
		6329	1			B	

TABLE 1.1

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO. OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83530	OV2-5	6351	1		X	B	
		7007	2			14.8	
		7011	1	1		B	
		7012	1			B	
		7013	2			B	
		7028		1			
		7032		1			
		7048	1			B	
		7049		1			
		7050		1			
		7063	4			B	
		7065	3			B	
		7066	3			B	
		7067	1	1		B	
		7068	2			B	
		7071	1			B	
		7073	1			B	
		7089	2	1		B	
		6294	1			B	
		6295	1			Flash, 11; Diffuse, 16	
		6304		1			
		6307	1	1		B	
		6308		1			
		6363		1			
		6364		1			
		7012		1			
		7075		1			
		7077		1			
		7089		1			

TABLE 1.j

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83533	ATS-3	6272	1			B	
		6273	3			B	
		6274	5			B	
		6275	3			B	
		6279	2			B	
		6280	7			B	
		6281	3			B	
		6282	3	2		B	
		6283	4			F	
		6294		1			
		6304		1			
		6307		1			
		6308		1			
		6314		1			
		6346	1				
		6363	1			B	
		7012	1			B	
		7013	1	2		B	
		7019	6			B	
		7028	2			B	
		7035		1			
		7049	1			B	
		7062	2			B	
		7063	4			B	
		7064	1			B	
		7065	3			B	
		7066	3			B	
		7068	1			B	
		7089		2			

TABLE 1.k

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83534	IS III F-6	6272		1			
		6273	1	1		B	
		6274		4			
		6275		2			
		6279		1			
		6294		1			
		6304		1			
		6307		1			
		6308		1			
		7049	2			F	
		7063	2			B	
		7064	1			B	
		7065	4			F	
		7066	2			B	
		7070	1			F	
		7077	1			B	
		7089		1			
83535	ATS-1	6272	1			B	
		6273	2	1		B	
		6274	3	1		B	
		6279	2			B	
		6280	9			B	
		6281	5			F	
		6282	4			B	
		6283	5			B	
		6294	1			B	
		6329	1			B	
		7050		1			
		7063	2			B	
		7065		3			

TABLE 1.1

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83546	IDCSP-14	7066		1			
		7069	1			B	
		7070	3			B	
		7071	1			B	
		7074	1			B	
		7078	1			B	
		7089		1			
		6279		1			
		6281	1	1		B	
		6282	4	4		13.5-F	
		6283	2	9		14-F	
		6294	1			B	
		6295	1		X	17.8	
		7025		1			
		7077		1			
83547	IDCSP-13	6273		1			
		6274		3			
		6287		1			
		6288		2			
		6290	1			16	
		6316		1			
		6330		1	X		
		7068		1			
83550	IDCSP-10	6279		1			
		6290	2			17	
		6316		1			
		6330		1	X		

TABLE 1.m

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83550	IDCSP-10	7025		2			
		7049		1			
		7050		1			
		7063		1			
		7074		1			
		7075	1			B	
		7077	2		X	13.5,17.2	
		7089		1			
83551	HATCH COVER	7050	1			F	
		7064	1			B	
		7069	3			F	
		7070	1			F	
		7071	1			B	
		7074	1			F	
		7075	1			B	
		7077	1			B	
		7078	1			B	
		7080		1			
		7089	1			F	
83558	RCA-A	6316	1			B	
		6329	1			B	
		6349	1			B	
		6351		1			
		7012	1			B	
		7014		1			
		7018		1			
		7019	1			B	
		7020	5			B	
		7027		1			

TABLE 1.n

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83560	SMS-2	7032	1			B	
		7035	1			B	
		7063		1			
		7064	1			B	
		7065	1	1		F	
		7066	1			B	
		7067	1			B	
		7071	1			B	
		7075	1			B	
		7077	1			B	
		7089	1			B	
		6275	2			B	
		6279	1			B	
		6280	4			B	
		6294	1			B	
		6304		1			
		6307	1	1		B	
		6308	1			B	
		6314		1			
		6315	1			B	
		6316	1			B	
		6325		1			
		6342		1			
		6343	1			B	
		6363		1			
		7019	8			B	
		7020	1			B	
		7025	1			B	
		7027	1			B	
		7028	2			B	
		7032		1			

TABLE 1.0

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO. OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m _v	
SOC NUMBER	COMMON NAME						
83567	SMS-1	7035		1			
		7040	2			14.6, 14.8	
		7049	1			B	
		7063	3			B	
		7064	1			B	
		7065	2			B	
		7066	2			B	
		7067	1			B	
		7070		1			
		7074		1			
		7078	1			B	
		7083		1			
		7089		1			
		6272	1	2		B	
		6273	3			B	
		6275	1	1		B	
		6279	1			B	
		6294	1			B	
		6304	1			B	
		6308	1			B	
		6314	1			B	
		6315	1			B	
		6316	1			B	
		6328	1			B	
		6330	12			B	
		6342		1			
		6343		1			
		6349	1			B	
		7004	1			B	
		7077	1			B	

TABLE 1.p

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO.OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83568	LES-6	7014		1	X		
		7019	16			B	
		7020	1			B	
		7025	1			B	
		7032	1			B	
		7035	1			B	
		7039	2			B	
		7040	1			14.1	
		7049	1			B	
		7063	5			B	
		7064		1			
		7065	3			B	
		7066	1	1		B	
		7073	1			B	
		7074	1			F	
		7075	2			F	
		7077	1			F	
		7078	3			B	
		7082		1			
		7089	1			B	
		6274		2			
		6279		1			
		6281	1			B	
		6282	3			B-F	
		6283	6	4		B	
		6290		3			
		7050	2	1		B	
		7063		1			
		7064	5			B	
		7075	2			B	
		7066	1			B	

TABLE 1.q

SATELLITE		UTD	NO. OF TIMES OBSER VED	NO. OF TIMES NOT FOUND	INTEGRA- TION USED	B, F or m v	
SOC NUMBER	COMMON NAME						
83569	WESTAR 2	7069	1			B	
		7070	1			B	
		7074		1			
		7075	1			B	
		7077	2			B	
		7078	2			B	
		7083	1			F	
		6280		7			
		6281		2			
		6294		1			
		6307		1			
		6308	1			B	
		6310		2			
		6315		1			
		6339		1			
		6343		1			
		6363		1			
		7012		1			
		7028	1			B	
		7029		1			
		7032	1			B	
		7035		1			
		7049	1			B	
		7063	1			B	
		7064	1			B	
		7065	4			B	
		7066	2			B	
		7067	1			B	
		7068		1			

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1976 UT0												
SATELLITE 270 . 1 . 260 . 1 . 290 . 1 . 300 . 1 . 310 . 1 . 320 . 1 . 330 . 1 . 340 . 1 . 350 . 1 . 360 . 1 .												
83501												
83502												
83503												
83505												
83507												
83509												
83512												
83513												
83523												
83530												
83533												
83534												
83535												
83546												
83547												
83550												
83558												
83560												
83567												
83568												
83569												
Sun DEC	-3.0°	-8.5°	-14.0°	-18.5°	-2.17°	-23.3°	-23.0°					
Moon Disk Visible	1.0	0.0	1.0	0.0	1.0	0.0	0.0					

Fig. 1a A graphical display of detections and non-detections of faint synchronous satellites at the GEODSS ETS from October, 1976 to March 1977. Sun declination and lunar phase are included for comparison.

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SATELLITE	1977 UT	0.1	10.1	20.1	30.1	40.1	50.1	60.1	70.1	80.1	90.1		
03501	0	0	0	X	0	0X	X	XXX	X	XXX	XX	X	
03502	00	0	0	X	X	X	X	XXX	X	X	XXX	XX	0
03503	0	0	0	0	0	0	0	0	0	0	0	0	
03505	0	0	XXX	X	X	0	0	0	0	0	0	0	
03507	X	0	0	0	0	0	0	0	0	0	0	0	
03509	6	XXX	0X	X	X	X	X	XXX	XX	X	X	X	
03512	0	0	0	0	0	0	0	0	0	0	0	0	
03513	0	X	0	0	0	0	0	0	0	0	0	0	
03523	X	XXX	0	0	0	0	0	0	0	0	0	0	
03530	0	0	0	0	0	0	0	0	0	0	0	0	
03533	XX	X	X	0	0	0	0	0	0	0	0	0	
03534	X	X	X	X	X	X	X	XXX	X	X	X	X	
03535	0	0	0	0	0	0	0	0	0	0	0	0	
03546	0	0	0	0	0	0	0	0	0	0	0	0	
03547	0	0	0	0	0	0	0	0	0	0	0	0	
03550	0	0	0	0	0	0	0	0	0	0	0	0	
03551	0	0	0	0	0	0	0	0	0	0	0	0	
03558	X	0	0XX	0	X	X	X	XXX	X	X	X	X	
03560	XX	X	XX	0	0	X	X	XXX	0	0	X	0	
03567	X	X	0	XX	X	X	X	XXX	XXX	XX	0	0	
03568	X	X	0	XX	X	X	X	XXX	XX	0X	XX	XX	
03569	0	0	0	X	0	X	X	XXX	XXX	XXX	XXX	XXX	
Sun DEC	-23.0°	-21.2°	-17.5°	-13.0°	-7.7°	-2.2°	+4.0°						
Moon Disk Visible	1.0	0.0	1.0	0.0	1.0	0.0	1.0	0.0	1.0	0.0	1.0	0.0	

Fig. 1b A graphical display of detections and non-detections of faint synchronous satellites at the GEODSS ETS from October, 1976 to March 1977. Sun declination and lunar phase are included for comparison.

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20. ABSTRACT (Continue on reverse side if necessary and identify by block number) This report presents a record of faint geosynchronous satellites detected at the GEODSS ETS from October 1976 to March 1977. The satellites reported have at one time or another been difficult or impossible to detect electro-optically because solar reflections from them were faint.		